## Higher-derivative massive gravity on a two-dimensional brane in 4D Minkowski space

Francisco A. Brito (UFCG)

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#### Introduction

- 1- Why low dimensional gravity?
- 2- Why higher derivative gravity?
- 3- Why massive gravity?
- 4- Why localization of gravity on a 2-brane?

#### Introduction

- 1- We shall see that these points are indeed connected with each other through a special mechanism of gravity localization called DGP (Dvali-Gabadadze-Porrati) scenario
- 2- Furthermore, the localization of gravity on a 2-brane is physically well-justified since we can see our real 3+1-dimensional universe as a stack of 2+1-dimensional branes

## **Gravity flutuations**

$$\eta_{\mu
u} o \eta_{\mu
u} + h_{\mu
u}$$

## RS scenario: zero mode

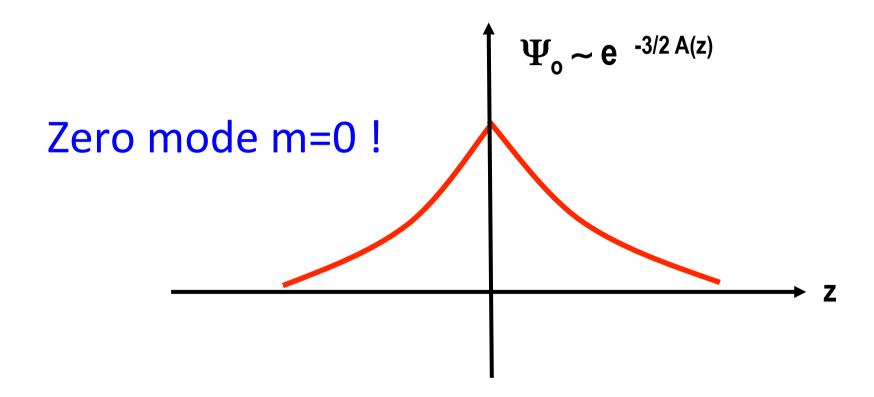
$$H\psi = m^2 \psi$$

$$H = Q^{\dagger} Q$$

$$Q = \partial_z + \frac{3}{2} \partial_z A(z)$$

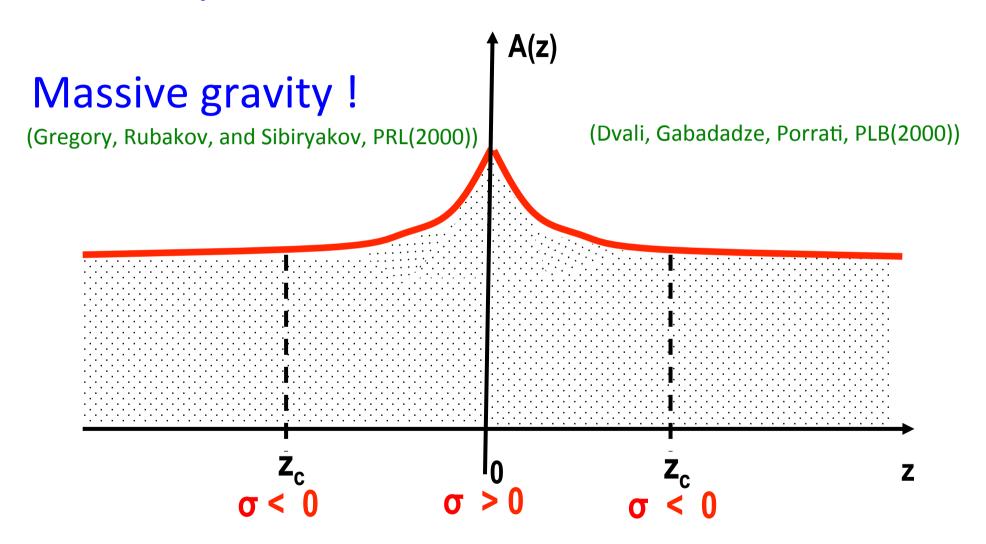
$$Q\psi_0 = 0 \longrightarrow \psi_0(z) = \exp\left(-\frac{3}{2} A(z)\right)$$

## RS scenario: finite volume



(Randall-Sundrum, PRL(1999))

## GRS/DGP scenario: infinite volume



# Newtonian law between two unit masses on the brane

$$V(r) = \frac{M_5^{-3}}{4rf(k)k} \left[ \sin\left(\frac{1}{2}\frac{f(k)r}{k}\right) \operatorname{Ci}\left(\frac{1}{2}\frac{f(k)r}{k}\right) - \cos\left(\frac{1}{2}\frac{f(k)r}{k}\right) \operatorname{Si}\left(\frac{1}{2}\frac{f(k)r}{k}\right) \right]$$

### The crossover scale

Brito, Losano, Fonseca, JCAP(2012)

$$r_c = k/f(k)$$
  $f(k) = \sqrt{\Lambda k^2 + 1}$ 

Large crossover for small cosmological const.

$$r_c = 1/\sqrt{\Lambda}$$
  $\Lambda k^2 >> 1$ 

Cosmological horizon!

## "">"Small" distance (r<<rc)

$$V(r) \sim \frac{M_5^{-3}\pi}{8rf(k)k} \sim M_5^{-3}\frac{r_c}{k^2r}$$

## Large distance (r>>rc)

$$V(r) \sim \frac{M_5^{-3}}{2 f(k)^2 r^2} \sim M_5^{-3} \frac{r_c^2}{k^2 r^2}$$

We inspire ourselves with the five-dimensional *DGP scenario* to extend it to an analogous set up in four-dimensions

$$S_{(D)} = \frac{M_D^2}{2} \int d^D x \sqrt{|g|} R^{(D)}$$

$$S_{(D-1)} = \frac{M_{D-1}}{2} \int d^{D-1}x \sqrt{|q|} R^{(D-1)}$$

$$q_{\mu\nu}(x) \equiv g_{\mu\nu}(x, w = 0)$$

$$m = \frac{M_{(D)}^2}{M_{(D-1)}} \equiv \frac{1}{r_c}$$

**General Framework** 

$$S = S_{(D)} + S_{(D-1)}$$

Dvali, Gabadadze & Porrati, PLB (2000)

Brito, Bazeia & Costa, PLB(2011)

Peng & Stojkovic, PRD(2014)

Brito, Bazeia & Costa, PLB (2015)

 $m=rac{M_{(D)}^2}{M_{(D-1)}}\equivrac{1}{r_c}$  q is the induced metric on the brane and rc is the 'crossover' scale

We inspire ourselves with the five-dimensional *DGP scenario* to extend it to an analogous set up in four-dimensions

$$S_{(4)} = \frac{M_4^2}{2} \int d^4x \sqrt{|g|} R^{(4)} \qquad S = S_{(4)} + S_{(3)}$$

$$S_{(3)} = \frac{M_3}{2} \int d^3x \sqrt{|q|} (\eta R^{(3)} + \gamma (R^{(3)})^2 + \delta R^{(3)}^{\mu\nu} R^{(3)}_{\mu\nu})$$

$$m = \frac{M_{(4)}^2}{M_{(3)}} \equiv \frac{1}{r_c}$$

Higher derivative terms!

The particle content of this theory is obtained by expanding the total action to linear order around the flat space:

$$g_{ab} = \eta_{ab} + H_{ab}$$
,  $q_{\mu 
u} = \eta_{\mu 
u} + h_{\mu 
u}, \quad N_{\mu} = n_{\mu}, \quad N = 1 + n.$   $H_{\mu 
u} = h_{\mu 
u}, \quad H_{\mu 4} = h_{\mu}, \quad H_{44} = 2n.$ 

After integrating out the bulk we obtain an effective 3D Lagrangian:  ${\cal L}_{Grav}=rac{1}{2}h_{\mu
u}O^{\mu
u,lphaeta}h_{lphaeta}$ 

In momentum space we have the operator:

$$O = \left[ \delta \frac{M_3}{4} k^4 + \eta \frac{M_3}{4} k^2 - M_3 m k \right] P^{(2)} - M_3 m k P^{(1)}$$

$$+\left[\left(\frac{8\gamma+3\delta}{4}\right)M_3k^4-\eta\frac{M_3}{4}k^2+M_3mk\right]P^{(0-s)}$$

$$+\sqrt{2}M_3mk(P^{(0-sw)}+P^{(0-ws)})$$

Brito, Bazeia & Costa, PLB (2015)

After integrating out the bulk we obtain an effective 3D Lagrangian:  ${\cal L}_{Grav}=rac{1}{2}h_{\mu
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The operator O is written in terms of the Barnes-Rivers operators  $P^{(2)}$ ,  $P^{(1)}$ ,  $P^{(0-s)}$ , ... which are spin projection operators that in three spacetime dimensions form a complete set and are defined as:

$$\begin{split} P_{\mu\nu,\kappa\lambda}^{(2)} &= \frac{1}{2} (\theta_{\mu\kappa}\theta_{\nu\lambda} + \theta_{\mu\lambda}\theta_{\nu\kappa} - \theta_{\mu\nu}\theta_{\kappa\lambda}), \\ P_{\mu\nu,\kappa\lambda}^{(1)} &= \frac{1}{2} (\theta_{\mu\kappa}\omega_{\nu\lambda} + \theta_{\mu\lambda}\omega_{\nu\kappa} + \theta_{\mu\nu}\omega_{\kappa\lambda}), \\ P_{\mu\nu,\kappa\lambda}^{(0-s)} &= \frac{1}{2} \theta_{\mu\nu}\theta_{\kappa\lambda}, \quad P_{\mu\nu,\kappa\lambda}^{(0-w)} &= \frac{1}{2} \omega_{\mu\nu}\omega_{\kappa\lambda}, \\ P_{\mu\nu,\kappa\lambda}^{(0-sw)} &= \frac{1}{\sqrt{2}} \theta_{\mu\nu}\omega_{\kappa\lambda}, \quad P_{\mu\nu,\kappa\lambda}^{(0-ws)} &= \frac{1}{\sqrt{2}} \omega_{\mu\nu}\theta_{\kappa\lambda}, \\ \theta_{\mu\kappa} &= \eta_{\mu\nu} - \frac{k_{\mu}k_{\nu}}{k^2}, \qquad \omega_{\mu\nu} &= \frac{k_{\mu}k_{\nu}}{k^2}, \end{split}$$

transverse and the longitudinal operators

$$O^{-1} = \frac{1}{\delta \frac{M_3}{4} k^4 + \eta \frac{M_3}{4} k^2 - M_3 mk} P^{(2)} - \frac{1}{M_3 mk} P^{(1)}$$

$$+\frac{\left(8\gamma+3\delta\right)M_3k^4-\eta M_3k^2-4M_3mk}{8M_3^2m^2k^2}P^{(0-w)}$$

$$+\frac{4\sqrt{2}M_3mk(P^{(0-sw)}+P^{(0-ws)})}{8M_3^2m^2k^2}.$$

The potential along the 2-brane is given in by

$$V(r) \sim \int \frac{d^3 \mathbf{k}}{2\pi^3} O^{-1} e^{i\mathbf{k}\cdot\mathbf{r}}$$

Taking only the spin-2 modes gives

$$V(r) \sim \int \frac{d^3 \mathbf{k}}{2\pi^3} \frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{\delta \frac{M_3}{4} \mathbf{k}^4 + \eta \frac{M_3}{4} \mathbf{k}^2 - M_3 m \mathbf{k}}$$

Here k<sup>2</sup> denotes the square of an Euclidean three-momentum.

#### The `cascading gravity'

According to intermediate scales of energy or (distance r) we may find the following cascading  $4D \rightarrow 3D \rightarrow 2D$  behavior:

$$V(r) \sim \int \frac{d^3 \mathbf{k}}{2\pi^3} \frac{e^{i\mathbf{k}\cdot\mathbf{r}}}{\delta \frac{M_3}{4} \mathbf{k}^4 + \eta \frac{M_3}{4} \mathbf{k}^2 - M_3 m \mathbf{k}}$$

$$V(r) \sim \frac{1}{r} \longrightarrow \ln(r) \longrightarrow r$$

(IR) k dominance  $\rightarrow$  k<sup>2</sup> dominance  $\rightarrow$  k<sup>4</sup> dominance (UV)

#### The `vanishing dimension'

This is also in the direction of the recently proposed "vanishing dimension" scenario where at high energy (or short scales) the physics appears to be lower dimensional!

Stojkovic, MPLA, (2013) Horava, PRD (2009)

$$V(r) \sim \frac{1}{r} \longrightarrow \ln(r) \longrightarrow r$$

(IR) k dominance  $\rightarrow$  k<sup>2</sup> dominance  $\rightarrow$  k<sup>4</sup> dominance (UV)

$$m^2 \to m^2(\square)$$
.

$$O^{-1} = \frac{1}{\delta \frac{M_3}{4} k^4 + \eta \frac{M_3}{4} k^2 - M_3 m^2(\square)} P^{(2)} - \frac{2}{2M_3 m^2(\square)} P^{(1)}$$

$$+\frac{(8\gamma+3\delta)\,M_3k^4-\eta M_3k^2-4M_3m^2(\Box)}{8M_3^2(m^2(\Box))^2}P^{(0-w)}$$

$$+\frac{4\sqrt{2}M_3m^2(\Box)(P^{(0-sw)}+P^{(0-ws)})}{8M_3^2(m^2(\Box))^2}.$$

$$m^2(\square) = L^{2(\alpha-1)}\square^{\alpha}$$

with L being a length scale and  $\alpha$  being a constant

$$m^2(\square) = L^{2(\alpha-1)}\square^{\alpha}$$

 $\alpha = 1/2$  recovers the DGP model

$$m^2(\Box) = mk \qquad (m=1/L)$$

$$m^2(\square) = L^{2(\alpha-1)}\square^{\alpha}$$

 $\alpha = 0$   $(m^2(\Box) = 1/L^2)$ ,  $\eta = -1$ ,  $\delta = 1/M^2$ ,  $\gamma = -3/8M^2$  gives the *Pauli-Fierz* mass term added to the *new massive* gravity in three dimensions Bergshoeff, Hohm & Townsend, PRL (2009)

$$O^{-1} = \frac{4}{M_3} \frac{1}{\left(\frac{1}{M^2}k^4 - k^2 - \frac{4}{L^2}\right)} P^{(2)} - \frac{L^2}{M_3} P^{(1)}$$

$$-\frac{L^4}{8M_3}\left(k^2+\frac{4}{L^2}\right)P^{(0-w)}+\frac{\sqrt{2}L^2}{2M_3}(P^{(0-sw)}+P^{(0-ws)}).$$

$$O^{-1} = \frac{4}{M_3} \frac{1}{\left(\frac{1}{M^2}k^4 - k^2 - \frac{4}{L^2}\right)} P^{(2)} - \frac{L^2}{M_3} P^{(1)}$$
$$-\frac{L^4}{8M_3} \left(k^2 + \frac{4}{L^2}\right) P^{(0-w)} + \frac{\sqrt{2}L^2}{2M_3} (P^{(0-sw)} + P^{(0-ws)}).$$

- There are massive poles in the sector of spin-2 graviton modes.
- One is the unitary mode of <u>positive mass</u> and <u>positive norm</u>.
- The other is the ghost of tachyonic mass and negative norm.

So both unitarity and causality are violated because of the existence of this tachyonic ghost!

$$m^2(\square) = L^{2(\alpha-1)}\square^{\alpha}, \quad \alpha \to 1 \qquad m^2(\square) \approx \square = k^2.$$

$$\begin{split} \delta &= 1/\mathsf{M}^2, \, \mathsf{\gamma} = -3/8 \mathsf{M}^2 \, \mathsf{gives} \\ O^{-1} &= \frac{4M^2}{M_3} \frac{1}{\left[k^2 - (4-\eta)M^2\right] k^2} P^{(2)} - \frac{1}{M_3 k^2} P^{(1)} \\ &- \frac{(4-\eta)}{M_2 k^2} P^{(0-w)} + \frac{\sqrt{2}}{2M_2 k^2} (P^{(0-sw)} + P^{(0-ws)}). \end{split}$$

The condition for the absence of tachyons is satisfied for both 'right' and 'wrong' η signs.

$$m^2(\square) = L^{2(\alpha-1)}\square^{\alpha}, \quad \alpha \to 1 \qquad m^2(\square) pprox \square = k^2.$$

$$\delta = 1/M^2$$
,  $\gamma = -3/8M^2$  gives

Bazeia, Brito & Costa, PLB (2015)

$$O^{-1} = \frac{4}{(4-\eta)M_3} \frac{m_2^2}{\left(k^2 - m_2^2\right)k^2} P^{(2)} - \frac{1}{M_3k^2} P^{(1)}$$

$$-\frac{(4-\eta)}{M_3k^2}P^{(0-w)}+\frac{\sqrt{2}}{2M_3k^2}(P^{(0-sw)}+P^{(0-ws)}).$$

Here  $m_2^2 = (4 - \eta)M^2$  is the massive spin-2 mode.

#### **Conclusions**

- 1- We consider the problem of localizing gravity in a 2-brane embedded in a 4D Minkowski space to address induction of high derivative massive gravity.
- 2- The structure of propagators shows a well-behaved higher-derivative massive gravity induced on the 2-brane.
- 3- We consider a special case in the generalized mass term of the graviton propagator, which ends up with a consistent higher order gravity.
- 4- The condition for the absence of tachyons is satisfied for both 'right' and 'wrong' signs of the Einstein-Hilbert term on the 2-brane.
- 5- By properly choosing the parameters of the theory one can find the Pauli-Fierz mass term added to the *new massive gravity* and recover the low dimensional DGP model.
- 6- It would be interesting to consider this study to pursue possible new aspects of the following cascading gravity:  $4D \rightarrow 3D \rightarrow 2D$

## Thanks a lot!





